Unconditional violation of the shot noise limit in photonic quantum metrology

Sergei Slussarenko,¹ Morgan M. Weston,¹ Helen M. Chrzanowski,^{1,2} Lynden K. Shalm,³ Varun B. Verma,³ Sae Woo Nam,³ and Geoff J. Pryde^{1,*}

¹*Centre for Quantum Dynamics, Griffith University, Brisbane, Queensland 4111, Australia*

²*Clarendon Laboratory, University of Oxford, Parks Road, Oxford OX1 3PU, UK*

³*National Institute of Standards and Technology, 325 Broadway, Boulder, Colorado 80305, USA*

Quantum metrology exploits quantum correlations to perform measurements with precision higher than can be achieved with classical approaches. Photonic approaches promise transformative advances in the family of interferometric phase measurement techniques, a vital toolset used to precisely determine quantities including distance, velocity, acceleration and various materials properties [1–3]. Without quantum enhancement, the precision limit in determining an unknown optical phase φ —i.e. the minimum uncertainty $\Delta\varphi$ —is the shot noise limit (SNL): $\Delta\varphi^{SNL} = 1/\sqrt{n}$, where n is the number of resources (e.g. photons) used. Entangled photons promise measurement sensitivity surpassing the shot noise limit achievable with classical probes. The maximally phase-sensitive state is a path-entangled state of definite number of photons N . Despite theoretical proposals stretching back decades [3, 4], no measurement using such photonic (definite photon number) states has unconditionally surpassed the shot noise limit: by contrast, all demonstrations have employed postselection to discount photon loss in the source, interferometer or detectors. Here, we use an ultra-high efficiency source and high efficiency superconducting photon detectors to respectively make and measure a two-photon instance of the maximally-phase-sensitive NOON state, and use it to perform unconditional phase sensing beyond the shot noise limit—that is, without artificially correcting for loss or any other source of imperfection. Our results enable quantum-enhanced phase measurements at low photon flux and open the door to the next generation of optical quantum metrology advances.

It has been known for several decades that probing with various optical quantum states can achieve *phase super-sensitivity*, i.e measurement of the phase with an uncertainty below the SNL [3, 4]. It has been shown theoretically that multi-photon entangled states, such as NOON states, may achieve super-sensitivity and can, in principle, saturate the Heisenberg limit (HL), the ultimate bound on sensitivity [3, 4, 7]. For this reason, they are of great interest for maximising the information that can be collected per photon, which is useful for investigating sensitive samples [8]. NOON states are superpositions of N photons across two arms of an interferometer, each of which is a single optical mode: $|\Psi_{NOON}\rangle = 1/\sqrt{2}(|N\rangle|0\rangle + |0\rangle|N\rangle)$. We use the term *photonic* to refer to states like this, because they possess definite photon number, and these photons are counted in detection. By contrast, we exclude from term “photonic” schemes using states of indefinite photon number and continuous wave-like measurement, such as squeezed states and homodyne detection. Such techniques have genuinely beaten the SNL, e.g. refs [5, 6], but work over narrow bandwidths and cannot directly achieve the theoretical maximal sensitivity per resource. The key feature of NOON and similar photonic states [17] is that they produce interference fringes that oscillate faster any classical interference pattern, a feature called *phase super-resolution* [9]. Super-resolution interference experiments have been reported using two-[10–13], three-[14], four-[15, 16], six- [9, 17] and eight- [18, 19] photon states.

Super-resolution, however, is not enough by itself to surpass the SNL [9, 20]: a high interference fringe visibility, and high transmission and detection efficiency are also required—they must exceed the threshold at which these imperfections cancel the quantum advantage. For imperfect NOON state interferometry, a handy estimate of this threshold was introduced in Ref. [9]: a genuine quantum advantage requires the interference visibility v and the combined single-photon transmission and detection efficiency η to satisfy $\eta^N v^2 N > 1$. (For precise evaluation of the potential for super-sensitivity, the Fisher information can be used to analyse experimental data, as described below.)

Here we performed the first phase sensing experiment with $N = 2$ photon NOON states that unconditionally demonstrates phase uncertainty below the SNL. This result was enabled by construction of a spontaneous parametric downconversion (SPDC) source [21], that was optimized for high photon transmission and required no spectral filtering in order to achieve high quantum interference visibility v . This allowed us to fully exploit the benefits of high-efficiency detector configuration [23], yielding an ultra-high heralding efficiency [22]—equivalent to the single-photon efficiency η . We emphasise that high-efficiency (95%) detection at 1550 nm wavelength has been available since 2008 [?], but photonic metrology beyond the SNL has not yet been achieved. This is because the capability to reach the required quantum interference and efficiency simultaneously requires a source capable of producing photons of exceptional spatial and spectral purity without filtering.

Unlike previous experiments, our measurement appa-

* g.pryde@griffith.edu.au

ratus does not require post-selection to achieve phase uncertainty below that achievable in an ideal, lossless classical interferometer. For our experimental apparatus, we expected $v \approx 0.98$ and symmetrical interferometer arm efficiencies $\eta \approx 0.8$ (which includes the detector efficiency), resulting in $\eta^N v^2 N \approx 1.23$. Thus, we anticipated a violation of the SNL, which we tested in two experiments described below.

For fair comparison with the SNL, an accurate accounting of resources is required. In the archetypal NOON-state phase sensing protocol, preparation and use of an N -photon NOON state constitutes a trial. In the ideal case, each trial leads to a detection event at the output of the interferometer. Since each trial gives only a little information about the phase, a number of such trials may be performed. In our work, two-photon NOON states were generated probabilistically at random times by the SPDC source. Each detection event (i.e. any combination of detector registrations) represented a *recorded* trial. We counted k such detection events to complete the protocol. However, due to imperfect transmission and detection efficiency η , some NOON states did not lead to detections. Furthermore, due to higher-order SPDC events (the occasional simultaneous emission of 4, 6, ... photons), the resources equivalent to multiple (2, 3, ...) trials were overlapped in time and could not be distinguished by our non-photon-number-resolving detectors. Therefore, the actual number of trials (i.e. the number of photon pairs passing the phase shift), $\tilde{k}$, was larger than the number of recorded trials. Because the ideal classical scheme is assumed to be lossless and to use all resources passing the phase shift, it must be attributed an effective number of resources $n = N\tilde{k} = 2\tilde{k}$. This makes the SNL harder to beat. For the loss and downconversion parameters of our experiment, the worst-case estimated, based on the lowest possible value for the overall experimental efficiency, we determined $\tilde{k}/k = 1.048125$.

Our experimental scheme is shown in Figure 1. We used collinear type-II parametric down-conversion, producing degenerate 1550 nm photon pairs [21]. Careful design and implementation of the source's output mode structure allowed us to achieve high fiber-coupling efficiency and state-of-the-art superconducting nanowire single photon detectors (SNSPDs) [23] provided high detection efficiency. The down-conversion process generates two photons in the $|1\rangle_H |1\rangle_V$ polarisation state ($H \equiv$ horizontal; $V \equiv$ vertical) in the same spatial mode, which can be written as the NOON-polarisation state $|\Psi\rangle = \frac{1}{\sqrt{2}}(|2\rangle_L |0\rangle_R + |0\rangle_L |2\rangle_R)$. These right- (R) and left-circular (L) polarisation modes constituted the two arms of the interferometer. A half-wave plate (HWP) set at an angle $\varphi/4$ relative to its optic axis was used to implement the birefringent phase shift φ between the arms. A common misconception about two-photon NOON states generated from SPDC is that the same phase sensitivity can be usefully achieved by using a pump photon (at half the wavelength) instead of the two-photon entangled state. However, this is clearly not correct for sensing in

Fig. 1. Experimental setup for the $N = 2$ NOON state optical interferometer. Laser pulses of ~ 170 fs duration, and centred at 775 nm, pump a 2 mm periodically poled KTP (pp-KTP) crystal, phase-matched for type-II collinear, group-velocity-matched, degenerate down-conversion at 1550 nm. A compensation crystal (1 mm KTP) was used to compensate the temporal walk-off between signal and idler photons. Pump and collection spatial modes were set by F (focusing) and C (collimating) lenses correspondingly. A silicon filter, AR-coated for 1550 nm, was used to block the pump beam. Transmitted (t) and reflected (r) modes of the interferometer, corresponding to H and V polarizations respectively, were separated by a polarizing beam-splitter (PBS) and then coupled into single mode fibres and sent to SNSPDs. A HWP mounted in an automated rotation stage is used to implement a controlled phase rotation, φ .

any material with dispersion.

After the phase shift, the modes were interfered on a polarising beam-splitter (PBS) and the output counting statistics were detected with SNSPDs and analyzed with coincidence or time-tag logic. The output signal consisted of three possible types of detection outcomes: C_{11} , a coincidence detection between both output modes; C_{20} , a detection occurring only in the transmitted output mode; and C_{02} , a detection occurring only in the reflected output mode. The numbers of each type of detection in a time period τ were, respectively, $c_{11}(\varphi)$, $c_{20}(\varphi)$ and $c_{02}(\varphi)$.

In order to test and calibrate our setup we first measured interference fringes. Detection events (≈ 250000 per phase value) were collected for a fixed amount of time for various $\varphi \in [0, 2\pi)$. We observed an interference visibility of $(98.9 \pm 0.02)\%$, calculated from fitting to the $c_{11}(\varphi)$ detection fringe. The transmissions of the reflected and transmitted outputs of the interferometer were measured to be $\eta_r = (79.41 \pm 0.09)\%$ and $\eta_t = (80.26 \pm 0.09)\%$, calculated from $c_{11}(0)/(c_{11}(0) + c_{20}(0))$ and $c_{11}(0)/(c_{11}(0) + c_{02}(0))$ ratios, respectively (a slight variation of transmission was observed when HWP was rotated, see Methods for details). Calculated transmissions include all the loss in the setup and the non-unit detection efficiency of SNSPDs. Probability fringes $p_{11}(\varphi)$, $p_{20}(\varphi)$ and $p_{02}(\varphi)$ were then obtained by fitting detection signals, $c_i(\varphi)$, $i \in \{11, 20, 02\}$, which were appropriately normalized for each phase value. We used the Fisher information per recorded trial, $\mathcal{F} = \sum_i \left(\frac{\partial \ln p_i}{\partial \varphi} \right)^2 p_i$ where $i \in \{11, 20, 02\}$, to quantify the phase sensitivity of our

Fig. 2. Experimentally measured output detection probability and the corresponding Fisher information. **a**, Yellow circles, blue triangles and red squares represent our experimentally determined $p_{11}(\varphi)$, $p_{20}(\varphi)$ and $p_{02}(\varphi)$ probabilities, respectively. Error bars are smaller than the markers. Lines represent a theoretical model with corresponding transmission and interference visibility parameters. **b**, The yellow curve represents the Fisher information per recorded trial determined from the probability fringes, $p_{11}(\varphi)$, $p_{20}(\varphi)$ and $p_{02}(\varphi)$, as a function of the unknown phase φ . The dashed red line represents the Fisher information (per recorded trial) at the SNL, while the solid blue line represents the SNL Fisher information taking into account the inefficiency and multi-photon emission—see text and Appendix for details. Shaded areas correspond to the 95% confidence region, derived from the uncertainty in the fit parameters.

phase measurement setup [3]. Our results (Fig. 2b) show a clear violation, for a range of phase values φ , of the adjusted SNL bound that takes into account the information in unrecorded trials: $\mathcal{F}_{SNL} = N\tilde{k}/k = 2.09625$.

For the second experiment, we performed phase sensing for individual settings of the phase shifter within the range where we expected to beat the SNL. At each setting, time-tag hardware was used to acquire detection events corresponding to $k = 10000$ trials. From the distribution of C_i events ($i \in \{11, 20, 02\}$), corresponding P_i probabilities were obtained by normalisation. The estimate of the phase, φ^{est} , was found by least squares fitting to the set of $p_i(\varphi)$ curves (Fig. 2a). The phase search range was restricted to $\varphi^{\text{est}} \in [0, \pi/2]$. This process was repeated for $s = 14520$ samples (for each angle of the

Fig. 3. Experimentally measured phase estimate and phase uncertainty. **a**, Lines represent theoretically modeled $p_{11}(\varphi)$, $p_{20}(\varphi)$ and $p_{02}(\varphi)$ probabilities from Fig. 2a. Yellow circles, blue triangles and red squares, correspond to average P_{11} , P_{20} , and P_{02} probabilities, correspondingly, calculated from 145200000 detection events per angle of HWP. Their positions on horizontal axis represent the measured phase estimate—see text for details. Error bars are smaller than the markers. **b**, Data points correspond to the standard errors of corresponding phase measurements of **a**. The purple line corresponds to standard error at the SNL, adjusted for correct number of resources $\tilde{N}$. The yellow line corresponds to the expected standard error, calculated from the Fisher information from Fig. 2b. The shaded areas correspond to 95% confidence regions, derived from the uncertainty in the fit parameters. The uncertainty in standard error was determined via the standard bootstrapping technique [24], similarly to Ref. [25].

HWP), determining φ_j^{est} for each sample j . The mean and standard error of $\{\varphi_j^{\text{est}}\}$ are shown in Fig. 3. This measurement procedure was repeated for a range of phase values around the region of interest. When compared to the standard error ($= 1/\sqrt{n^{\text{tot}}}$) that is achievable with $n^{\text{tot}} = ns = N\tilde{k}s = 304375500$ classical resources (adjusted for loss and higher order terms, as before) per data point, our results show a clear advantage of our quantum approach.

In conclusion, we demonstrated unconditional violation of the SNL in photonic phase sensing. We recover the estimate of the phase shift applied to the mode, and its corresponding standard error, directly from our measure-

ment data without additional adjustments. Moreover, all the parameters necessary for the calculation of the SNL of our measurement apparatus, such as circuit loss, pair and multi-pair generation probability, were calculated directly from our measurement data. Our results are not only of fundamental interest, but are also directly applicable to a phase measurement scenario where low photon flux is required, such as measurement of light-sensitive materials [8]. We anticipate that our technique can be extended, using loss-tolerant approaches [26] and number-resolving detectors [27], to perform sub-SNL sensing with significantly higher photon numbers in the near future.

I. APPENDIX

A. Photon source and characterization

High heralding efficiency (Klyshko efficiency [22]) and high interference visibility were essential to the success of this demonstration. These features were achieved by using a spontaneous parametric down-conversion (SPDC) source that operated at the group velocity matching conditions to generate frequency uncorrelated photon pairs, removing the need for spectral filtering that is typically required in conventional SPDC sources. The heralded single photon source was pumped by a mode-locked Ti:Sapphire laser with 81 MHz repetition rate, 775 nm wavelength and 6 nm FWHM bandwidth. SPDC from a nonlinear periodically poled KTP crystal (pp-KTP, poling period 46.20 μm) phase matched for type-II collinear operation produced degenerate photon pairs centred at 1550 nm wavelength and with ≈ 15 nm FWHM bandwidth. We used a pump waist size of 170 μm and a signal and idler collection waist size of 50 μm [21]. Together with the high efficiency SNSPDs [23], this allowed us to achieve symmetric heralding efficiencies of $(82 \pm 2)\%$. An additional 1 mm KTP crystal, with optic axis rotated by 90° with respect to the optic axis of pp-KTP, was placed after the downconverter to compensate for the temporal walk-off (due to birefringence) in the 2 mm pp-KTP crystal. Non-classical interference visibilities between the signal and idler photons were measured to be $(98.9 \pm 0.2)\%$, with no background subtraction.

B. Detection

There is one superconducting nanowire single photon detector (SNSPD) coupled to each of the interferometer outputs. We record three types of events: C_{11} , a coincidence detection between both output modes; C_{20} , a detection occurring only in the transmitted output mode; and C_{02} , a detection occurring only in the reflected output mode. Since the SNSPDs cannot resolve photon number, the latter two events may arise from: the case where both photons in the two-photon NOON state went to the same detector; a case where one photon was lost

and therefore only one photon is recorded; or the case of a dark count. We do not artificially remove signals from these latter two cases. We note that the rate of dark counts is small compared to the rate of real events.

C. Resource counting

The total transmission (including detection efficiency) is easily determined when the phase shift is set to zero. However, we have also observed that the circuit transmission can slightly vary with the rotation of the HWP that implements the phase shift. In order to verify transmission at every value of φ deviates only slightly from this, we build a theoretical model of the interferometer, which includes an imperfect interference on the PBS and loss in two of the output modes. Then, at each rotation angle of the HWP we find transmission of both output modes ($\eta_r(\varphi)$ and $\eta_t(\varphi)$) by finding the closest theoretical match to the measured data with least squares minimization. We confirm that the transmission variation stays below $\approx 1\%$. Moreover, the standard deviation in total number of detection events per point in Fig. 2, which were acquired over a fixed amount of time, was measured to be $\approx 0.2\%$, confirming that the variation of our setup parameters was very small.

In order to estimate the true number of resources we use, we calculated the probability of having at least one photon transmitted through the system. For the three possible outcomes, this is given by

$$\begin{aligned}\eta_{11}(\varphi) &= 1 - (1 - \eta_t(\varphi))(1 - \eta_r(\varphi)), \\ \eta_{20}(\varphi) &= 1 - (1 - \eta_t(\varphi))(1 - \eta_t(\varphi)), \\ \eta_{02}(\varphi) &= 1 - (1 - \eta_r(\varphi))(1 - \eta_r(\varphi)).\end{aligned}\tag{1}$$

The lowest efficiency $\eta_{\min} = \min_{\varphi,j}(\eta_j(\varphi))$, for $j \in \{11, 20, 02\}$ and $\varphi \in [0, 2\pi)$, observed over entire range of φ out of the three outcomes was $\eta_{\min} \approx 95.56\%$. We chose the lowest efficiency $\eta_{\min}(\varphi) = \min_j(\eta_j(\varphi))$, calculated at each angle of the HWP (i.e. at each φ) in order to calculate the number of trials accordingly. This results in an overestimate of the number of photons used in our experiment—and therefore used to calculate the effective SNL—setting a higher SNL bound than was actually the case, and making it harder to violate the SNL. The uncertainty in the estimated transmission and the SNL bound was calculated by error propagation of the uncertainty in c_i through the theoretical model of the interferometer. We have observed very strong dependence of the estimated parameters on $c_{11}(\varphi)$ near the $\varphi = \pi/2 + k\pi$, $k \in \mathbb{Z}$, resulting in an increased parameter uncertainty in those regions.

Another important source of additional resources that are not directly detected in the protocol is the possible emission of multiple photon pairs in the SPDC process. For the pump power of ≈ 100 mW used in our experiment, we measured this probability to be $\xi = (0.1550 \pm 0.0004)\%$, relative to the probability of generating a single pair. This was done by comparing the

number of photon detection events within a fixed time window with the actual number of laser pulses in that window, acquired with a separate trigger signal. We determined ξ by carefully determining the single pair generation probability, with loss taken into account.

The actual number of trials (pairs of photons generated, for $N = 2$) $\tilde{k}$ is related to the number of recorded

trials k by

$$\tilde{k}(\varphi) = \frac{k(1 + \xi)}{\eta_{\min}(\varphi)}. \quad (2)$$

Equivalently, the actual average number of resources per recorded trial is

$$\frac{N\tilde{k}}{k} = \frac{2(1 + \xi)}{\eta_{\min}(\varphi)}, \quad (3)$$

for $N = 2$.

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- [1] C. M. Caves, *Phys. Rev. D* **23**, 1693 (1981).
 - [2] H. M. Wiseman and G. J. Milburn, *Quantum Measurement and Control* (Cambridge University Press, Cambridge, England, 2009).
 - [3] V. Giovannetti, S. Lloyd, and L. Maccone, *Nature Photon.* **5**, 222 (2011).
 - [4] R. Demkowicz-Dobrzański, M. Jarzyna, and J. Kołodyński, *Prog. Opt.* **60**, 345 (2015).
 - [5] H. Yonezawa, D. Nakane, T. A. Wheatley, K. Iwasawa, S. Takeda, H. Arao, K. Ohki, K. Tsumura, D. W. Berry, T. C. Ralph, H. M. Wiseman, E. H. Huntington, and A. Furusawa, *Science* **337**, 1514 (2012).
 - [6] J. Aasi, J. Abadie, B. Abbott, R. Abbott, T. Abbott, M. Abernathy, C. Adams, T. Adams, P. Addesso, R. Adhikari, *et al.*, *Nature Photon.* **7**, 613 (2013).
 - [7] J. P. Dowling, *Contemp. Phys.* **49**, 125 (2008).
 - [8] F. Wolfgramm, C. Vitelli, F. A. Beduini, N. Godbout, and M. W. Mitchell, *Nature Photon.* **7**, 28 (2013).
 - [9] K. J. Resch, K. L. Pagnell, R. Prevedel, A. Gilchrist, G. J. Pryde, J. L. O'Brien, and A. G. White, *Phys. Rev. Lett.* **98**, 223601 (2007).
 - [10] Z. Y. Ou, X. Y. Zou, L. J. Wang, and L. Mandel, *Phys. Rev. A* **42**, 2957 (1990).
 - [11] J. G. Rarity, P. R. Tapster, E. Jakeman, T. Larchuk, R. A. Campos, M. C. Teich, and B. E. A. Saleh, *Phys. Rev. Lett.* **65**, 1348 (1990).
 - [12] E. J. S. Fonseca, C. H. Monken, and S. Pádua, *Phys. Rev. Lett.* **82**, 2868 (1999).
 - [13] H. S. Eisenberg, J. F. Hodelin, G. Khoury, and D. Bouwmeester, *Phys. Rev. Lett.* **94**, 090502 (2005).
 - [14] M. W. Mitchell, J. S. Lundeen, and A. M. Steinberg, *Nature* **429**, 161 (2004).
 - [15] P. Walther, J.-W. Pan, M. Aspelmeyer, R. Ursin, S. Gasparoni, and A. Zeilinger, *Nature* **429**, 158 (2004).
 - [16] T. Nagata, R. Okamoto, J. L. O'Brien, K. Sasaki, and S. Takeuchi, *Science* **316**, 726 (2007).
 - [17] G. Xiang, H. Hofmann, and G. J. Pryde, *Sci. Rep.* **3** (2013).
 - [18] W.-B. Gao, C.-Y. Lu, X.-C. Yao, P. Xu, O. Gühne, A. Goebel, Y.-A. Chen, C.-Z. Peng, Z.-B. Chen, and J.-W. Pan, *Nature Phys.* **6**, 331 (2010).
 - [19] X.-L. Wang, L.-K. Chen, W. Li, H.-L. Huang, C. Liu, C. Chen, Y.-H. Luo, Z.-E. Su, D. Wu, Z.-D. Li, H. Lu, Y. Hu, X. Jiang, C.-Z. Peng, L. Li, N.-L. Liu, Y.-A. Chen, C.-Y. Lu, and J.-W. Pan, *Phys. Rev. Lett.* **117**, 210502 (2016).
 - [20] R. Okamoto, H. F. Hofmann, T. Nagata, J. L. O'Brien, K. Sasaki, and S. Takeuchi, *New J. Phys.* **10**, 073033 (2008).
 - [21] M. M. Weston, H. M. Chrzanowski, S. Wollmann, A. Boston, J. Ho, L. K. Shalm, V. B. Verma, M. S. Allman, S. W. Nam, R. B. Patel, S. Slussarenko, and G. J. Pryde, *Opt. Express* **24**, 10869 (2016).
 - [22] D. N. Klyshko, *Sov. J. Quantum Electron.* **10**, 1112 (1980).
 - [23] F. Marsili, V. B. Verma, J. A. Stern, S. Harrington, A. E. Lita, T. Gerrits, I. Vayshenker, B. Baek, M. D. Shaw, R. P. Mirin, and S. W. Nam, *Nature Photon.* **7**, 210 (2013).
 - [24] A. C. Davison and D. V. Hinkley, *Bootstrap methods and their application*, Vol. 1 (Cambridge university press, 1997).
 - [25] B. L. Higgins, D. W. Berry, S. D. Bartlett, H. M. Wiseman, and G. J. Pryde, *Nature* **450**, 393 (2007).
 - [26] J. C. F. Matthews, X.-Q. Zhou, H. Cable, P. J. Shadbolt, D. J. Saunders, G. A. Durkin, G. J. Pryde, and J. L. O'Brien, *NPJ Quantum Inf.* **2**, 16023 (2016).
 - [27] G. Harder, T. J. Bartley, A. E. Lita, S. W. Nam, T. Gerrits, and C. Silberhorn, *Phys. Rev. Lett.* **116**, 143601 (2016).